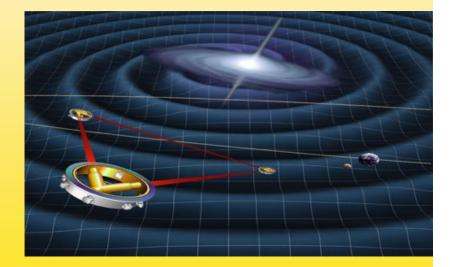


The Warped Side of the Universe Kip Thorne



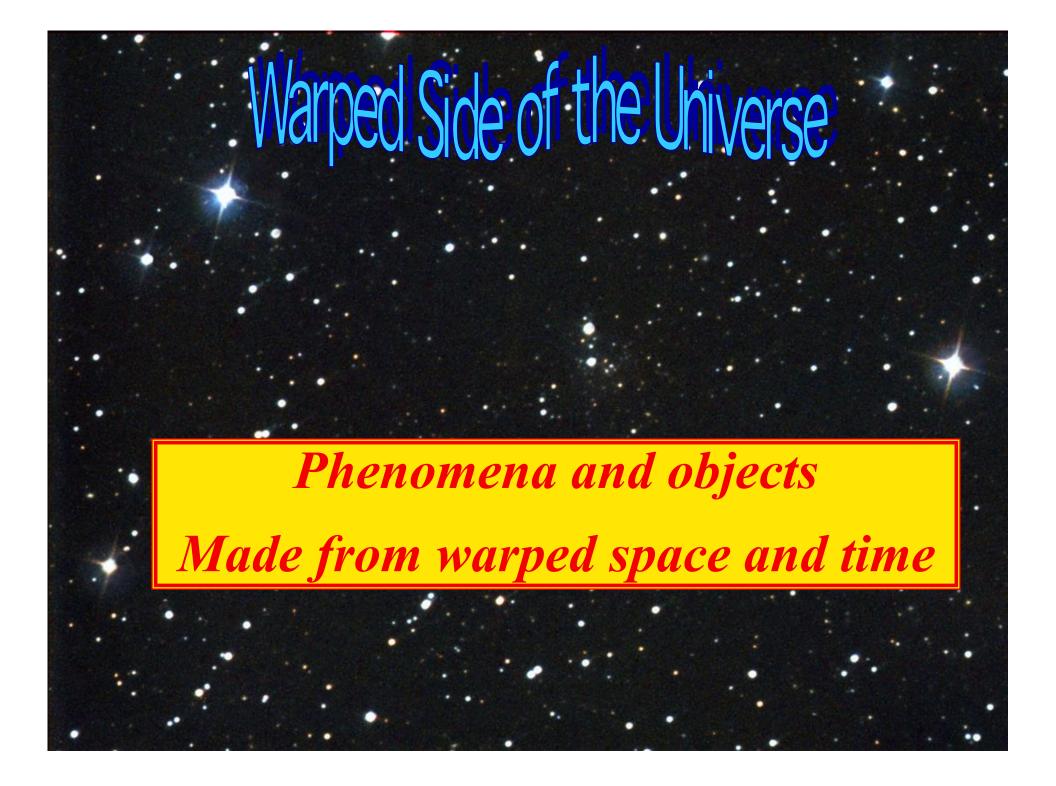


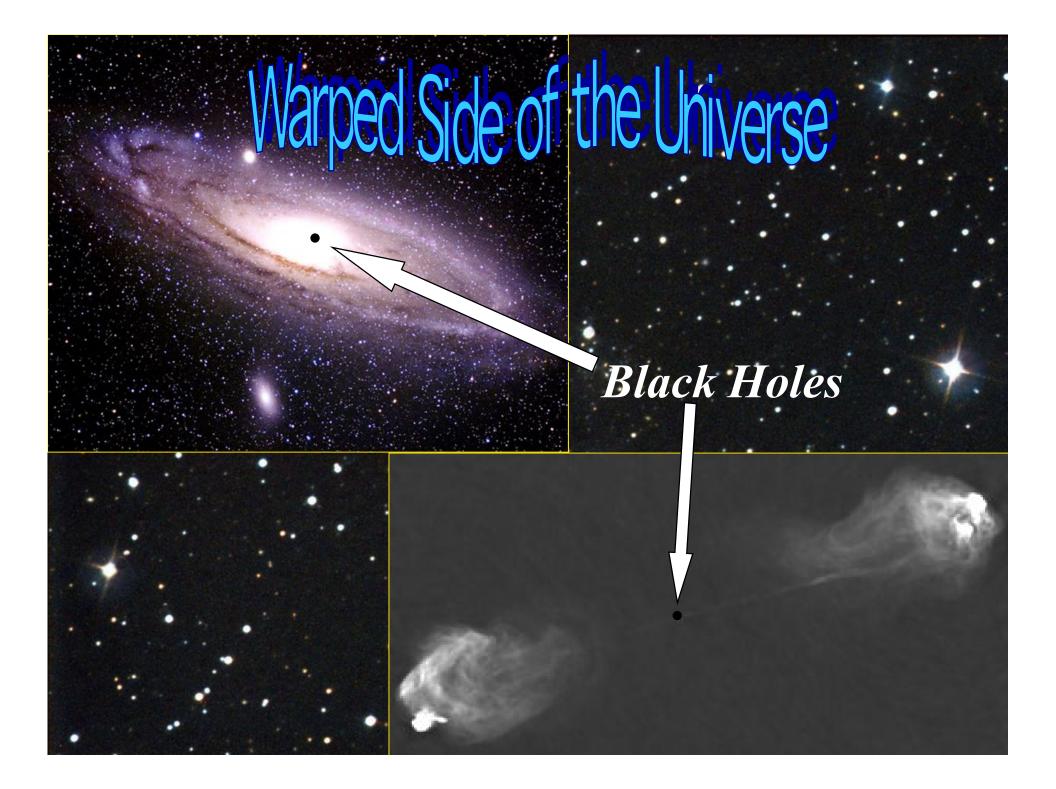
Colloquium Ehrenfestii, University of Leiden, 16 Sept 2009

Lorentz Lectures: Gravitational Waves

- Understandable without prior knowledge of general relativity
- But knowledge of general relativity will help
- Slides available (pdf) late Thursday nights at

http://www.its.caltech.edu/~kip/LorentzLectures/





Black Hole!s Spacetime Geometry

Curvature of Space

- Rotational Motion of Space
- Warping of Time

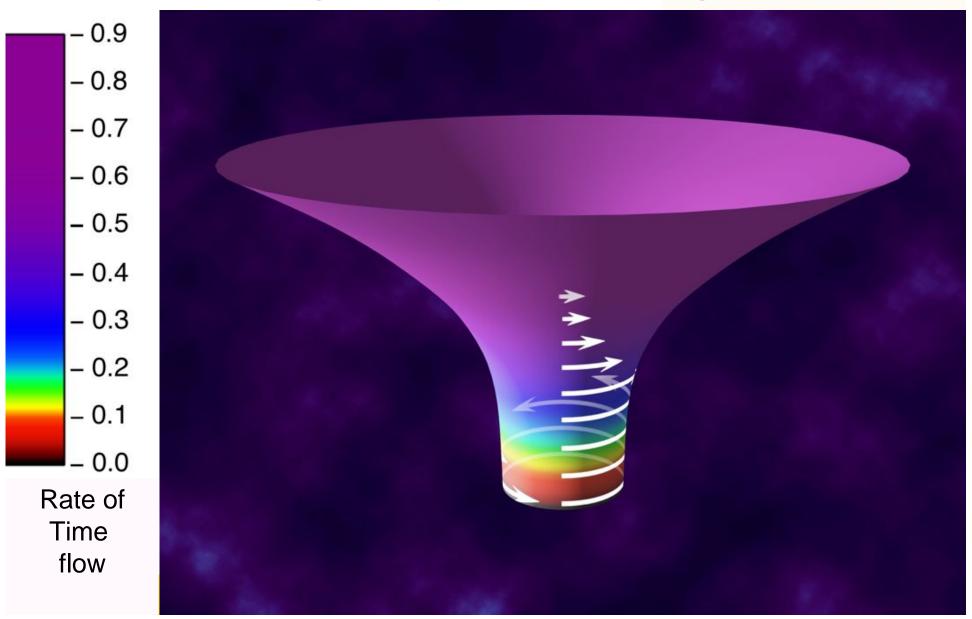
e

$$g_{\theta\theta} d\theta^2 + g_{\phi\phi} (d\phi - \omega dt)^2 - \alpha^2 dt^2$$

 $ds^{2} = g_{rr}dr^{2} + g_{\theta\theta} d\theta^{2} + g_{\phi\phi} (d\phi - \omega dt)^{2} - \alpha^{2} dt^{2}$ space curvature space rotation time warp *3-metric* shift function *lapse* function

Kerr Metric

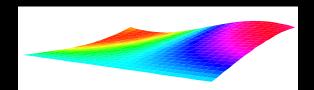
Map of spacetime geometry for fast spinning hole a/M=0.998

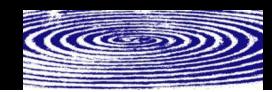


• The Big Bang Singularity

- Our Universe as a Brane in a higher-dimensional bulk
- Cosmic String
- Singularity inside a black hole
- Naked Singularity
- Wormhole
- Gravitational Waves











 $C/R = 2\pi (1 - 4G\mu/c^2)$

Warped Side of the Universe OTHER EXAMPLES

WHICH ARE REAL?

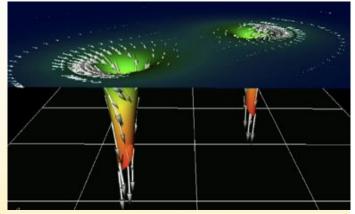
WHAT ELSE?

Probing the Warped Side: Tools

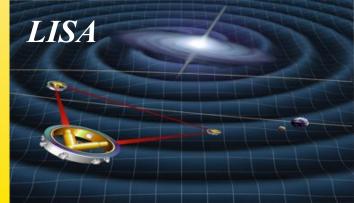
- What kinds of objects *might* exist?
 » General Relativity Theory

 Progress has slowed...
 - » Numerical Relativity
 - Exciting new era...
 - Part 1 of Colloquium
- What kinds of objects *do* exist?
 » Electromagnetic observations

 –Limited information
 - » Gravitational-Wave observations
 - -Ideal tool for probing the
 - . Warped Side
 - -Part 2 of Colloquium

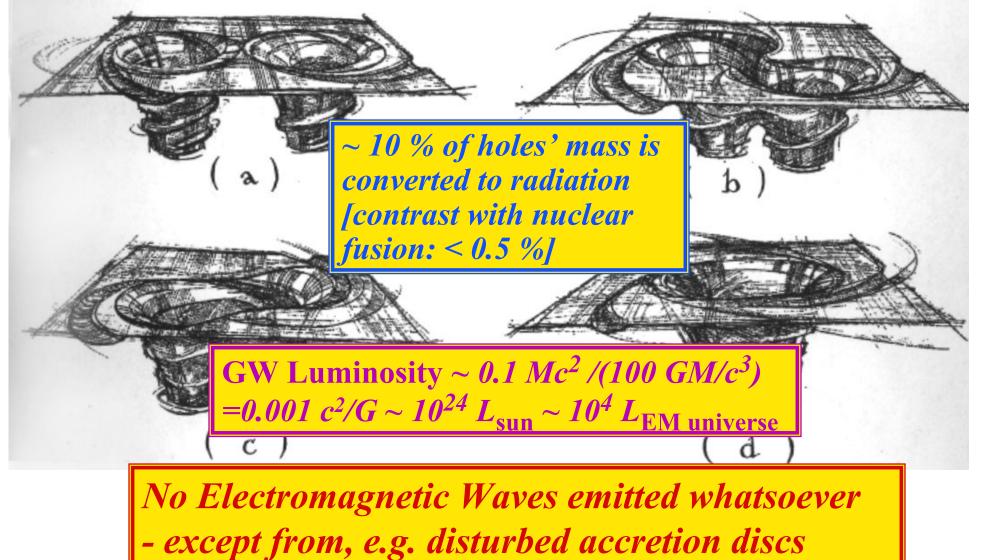




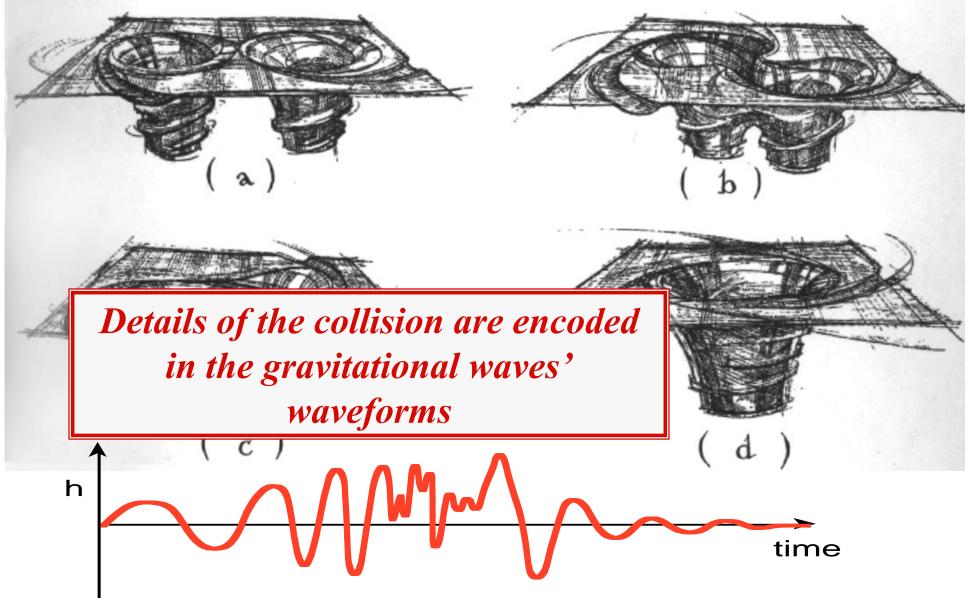


Part 1 Numerical Relativity

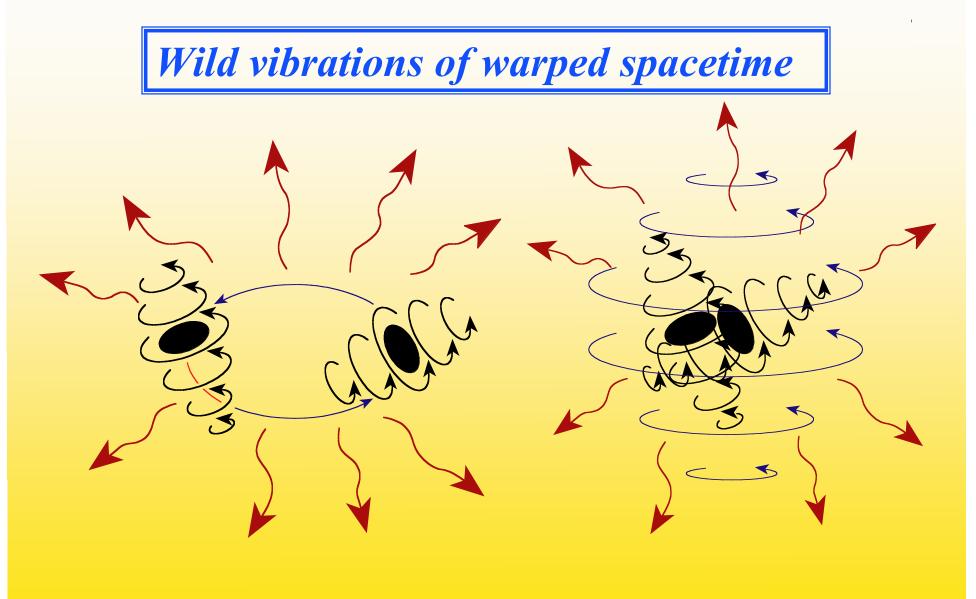
The "Holy Grail": Collisions of Black Holes - The most violent events in the Universe



Collisions of Black Holes: The most violent events in the Universe

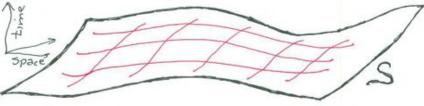


Why are Black-Hole Collisions Interesting?



Numerical Relativity: How is it Done?

- Evolve the geometry of spacetime not fields in spacetime
- Choose an initial spacelike 3-dimensional surface S
 - » Put a coordinates on S



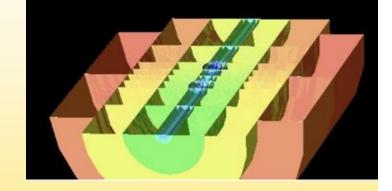
- Specify: 3-metric g_{ij} and Extrinsic Curvature K_{ij} of S
 - » Subject to constraint equations [analogues of Div B = 0]
- Lay out coordinates to future by specifying Lapse function α and Shift function β^{i}
- Integrate 3-metric forward in time via dynamical equations
- Build 4-metric of spacetime

 $ds^2 = -\alpha^2 dt^2 + g_{ij} (dx^i - \beta^i dt) (dx^j - \beta^j dt)$

Numerical Relativity

Two Mature Approaches

- Finite-difference
 - » Robust, power-law convergence [Astrophysics]
- Spectral
 - » More complicated, less robust
 - but exponential convergence
- » Fast; high accuracy [GWs] Two Major Pitfalls



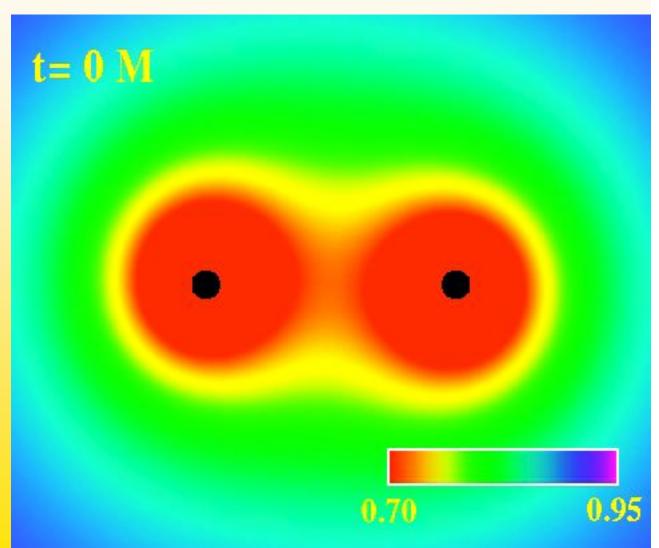
- Constraint-Violation Instabilities:
 - » Slight initial error in constraints (analog of Div $\mathbf{B} = 0$) blows up in time
 - » Solved in 2005 after ~ 5 years of struggle
- Coordinates become singular
 - » Only now becoming robustly solved in spectral code Szilagyi, Lindblom, Scheel, PRD & arXiv - submit this weekend

Numerical Relativity Breakthrough

• The first successful simulations, May 2005: *Frans Pretorius* (then at Caltech; now Princeton) finite-difference

Identical holes, not spinning

Lapse Function



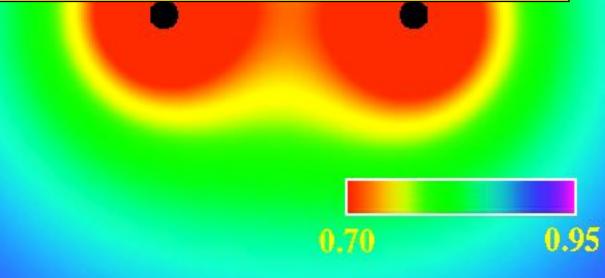
Numerical Relativity Breakthrough

The first successful simulations, May 2005: Frans Pretorius (then at Caltech; now Princeton) finite-difference

t = 0 M

Ide

- Followed, 6 months later, by finite-difference success atU Texas Brownsville (Campanelli, Lousto, Zlochower) not
 - Goddard Spaceflight Center (Baker, Centrella, Choi, ...) Then many others



Lapse Function

Numerical Relativity Groups Today	
 Princeton [Pretorius] University of British Columbia [Choptuik] 	
 University of Illinois, Urbana [Shapiro] 	
 University of Chicago [Khokhlov] 	
 University of Texas, Austin [Matzner] 	
 Louisiana State University [Seidel, Pullin, …] 	
 Goddard Spaceflight Center [Centrella, Baker] 	Finite Difference
 Rochester Institute of Technology [Campanelli, Lousto] 	I intic Dijjerenee
Oakland University [Garfinkle]	
Florida Atlantic [Miller, Tichy]	
Albert Einstein Institute [Rezzolla]	
 University of Jena [Bruegmann] 	
 University of Tokyo [Shibata] 	
 Cornell/Caltech [Teukolsky, Lindblom, Kidder, Scheel, F 	Pfeiffer] SpEC
New Groups this year & last -	Spectral Einstein Code)
» Georgia Tech [Laguna, Shoemaker]	Difference
» Perimeter Institute / U. Guelph [Lehner]	Dijjerence
» Canadian Institute for Theoretical Astrophysics [Pfeiffer] SpEC	
» U. Maryland [Tiglio] ← <i>SpEC</i>	18

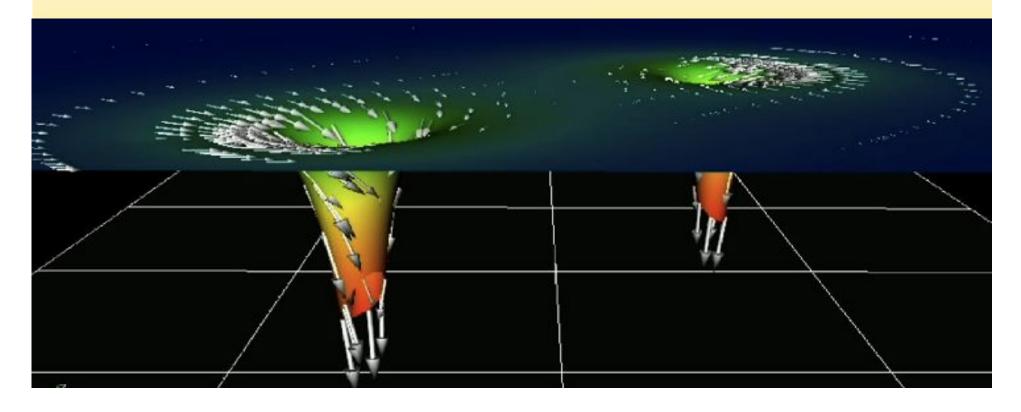
Numerical Relativity Groups Today

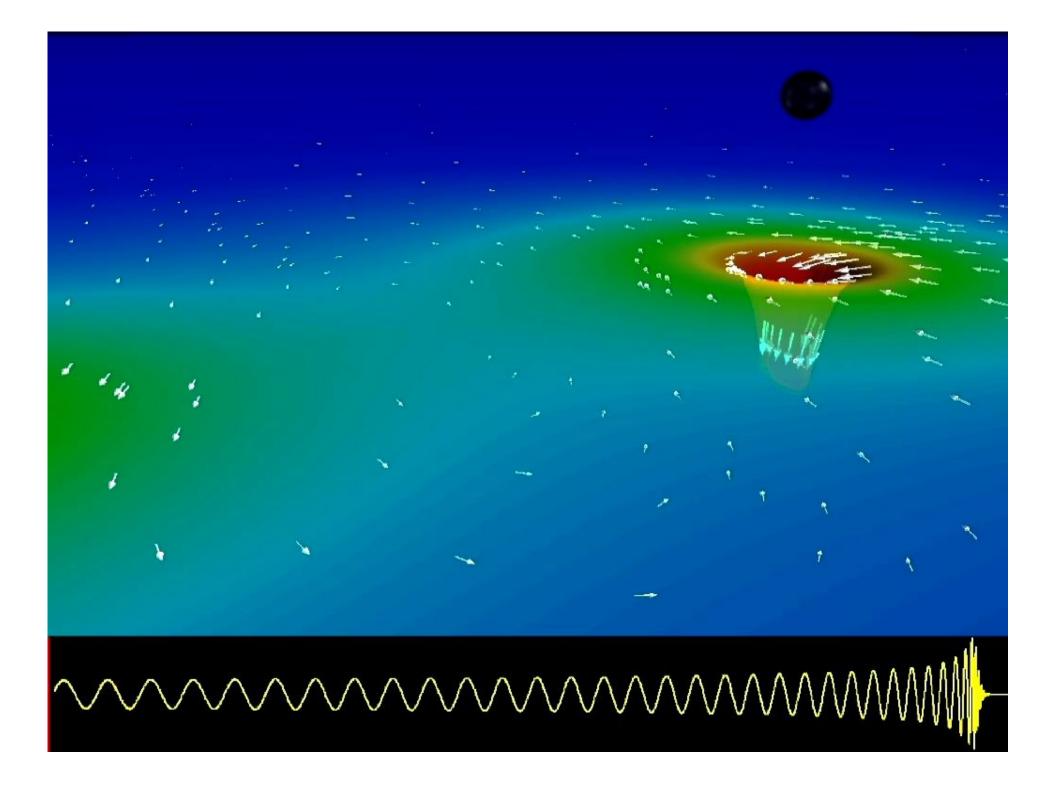
- Princeton [Pretorius]
- University of British Columbia [Choptuik]
- University of Illinois. Urbana [Shapiro]
- 2005 2009:
- Finite Difference groups: astrophysical studies at
- moderate accuracy
 - SpEC: solve coordinate problems make code robust
 - . embark on GW studies at high accuracy, fast speed;
 - . begin exploring nonlinear dynamics of warped spacetime
- University of Jena [Bruegmann]
- University of Tokyo [Shibata]
- New Groups this year & last -
 - » Georgia Tech [Laguna, Shoemaker]
 - » Perimeter Institute / U. Guelph [Lehner]
 - » Canadian Institute for Theoretical Astrophysics [Pfeiffer] SpEC
 - » U. Maryland [Tiglio] ← SpEC

(Spectral Einstein Code)

State of the Art Today. SpEC Example:

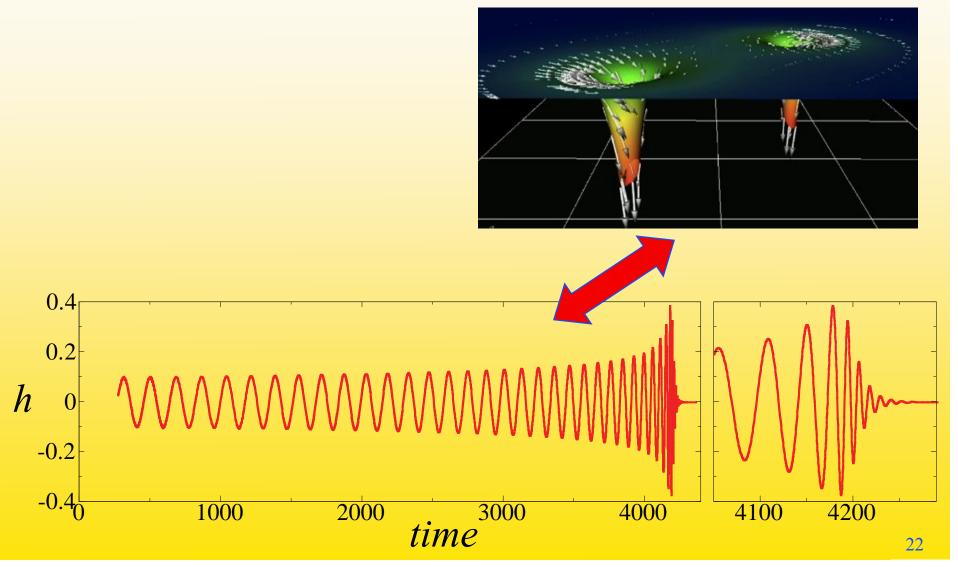
- Identical holes, not spinning
- 16 orbits, collision, merger, and ringdown
- Gravitational waveforms cumulative phase error
 ~ 0.01 radians Cornell/Caltech: Kidder, Lindblom Pfeiffer, Scheel, Teukolsky...





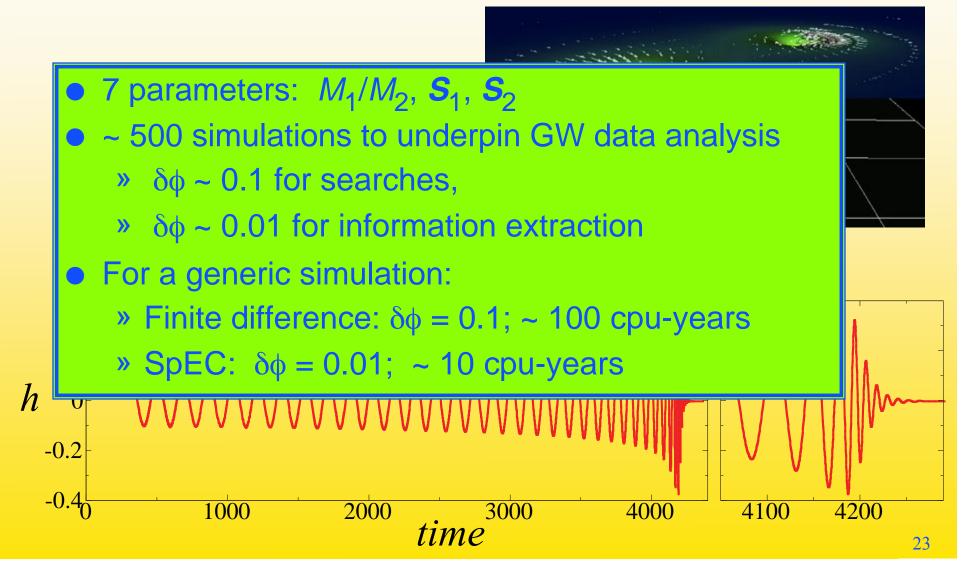
"Rosetta Stone"

Numerical simulations and theory provide "rosetta stone" for interpreting observed Gravitational Waveforms



"Rosetta Stone"

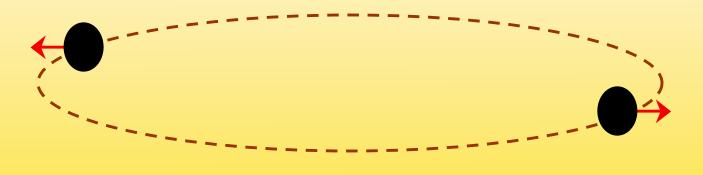
Numerical simulations and theory provide "rosetta stone" for interpreting observed Gravitational Waveforms



Nonlinear Dynamics of Warped Spacetime

• Spinning Holes

- » Rochester Institute of Technoogy: Campanelli, Lousto, Zlochower
- » Finite-difference techniques



- Simulation: Manuela Campanlli Carlos Lousto Yosef Zlochower
- Visualization: Hans-Peter Bischof
- CCRG RIT

Copyright - CCRG - 2009

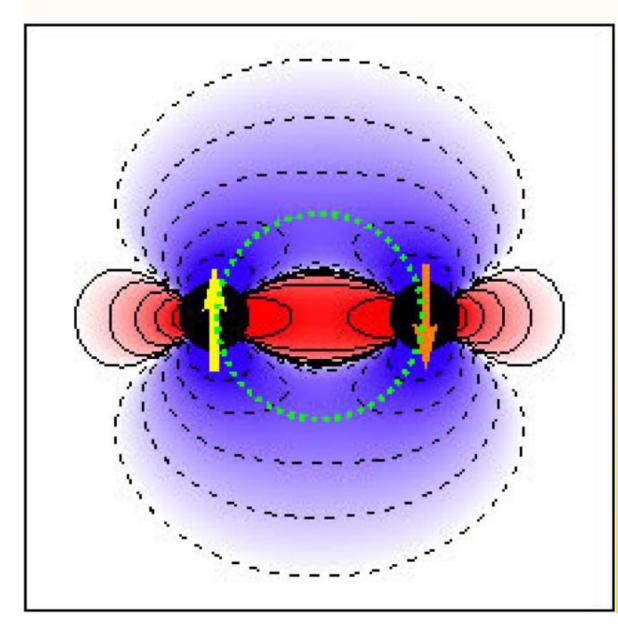
Analogous to 2 Vortices in a Fluid

from *Vorticity* by Asher H. Shapiro (National Comittee on Fluid Mechanics Films, ca 1960)



Nonlinear Dynamics of Warped Spacetime

Explanation (Pretorius): "Frame Dragging" + Spin/Curvature Coupling

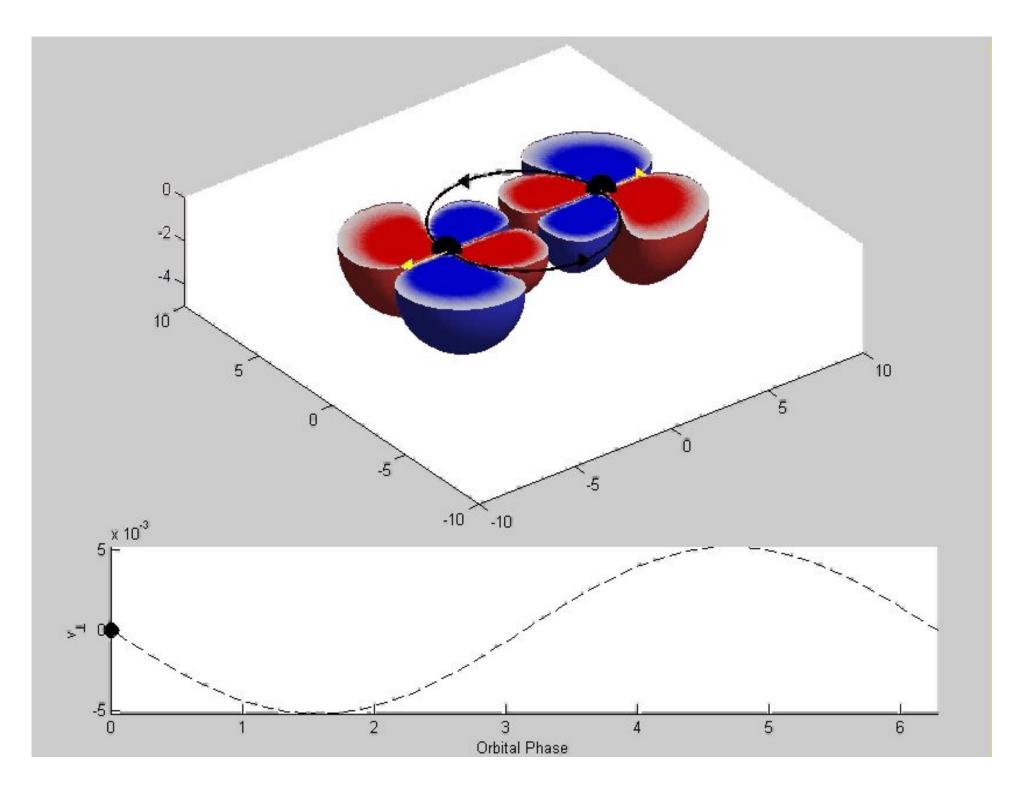


Momentum conservation: Chen, Keppel, Nichols, Kip

Field momentum:

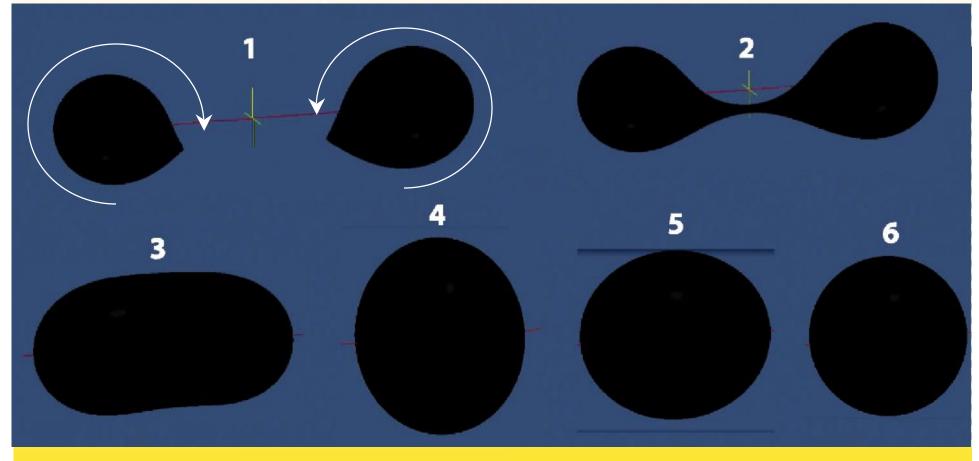
$$-\frac{1}{4\pi}\vec{g}\times\vec{H}$$

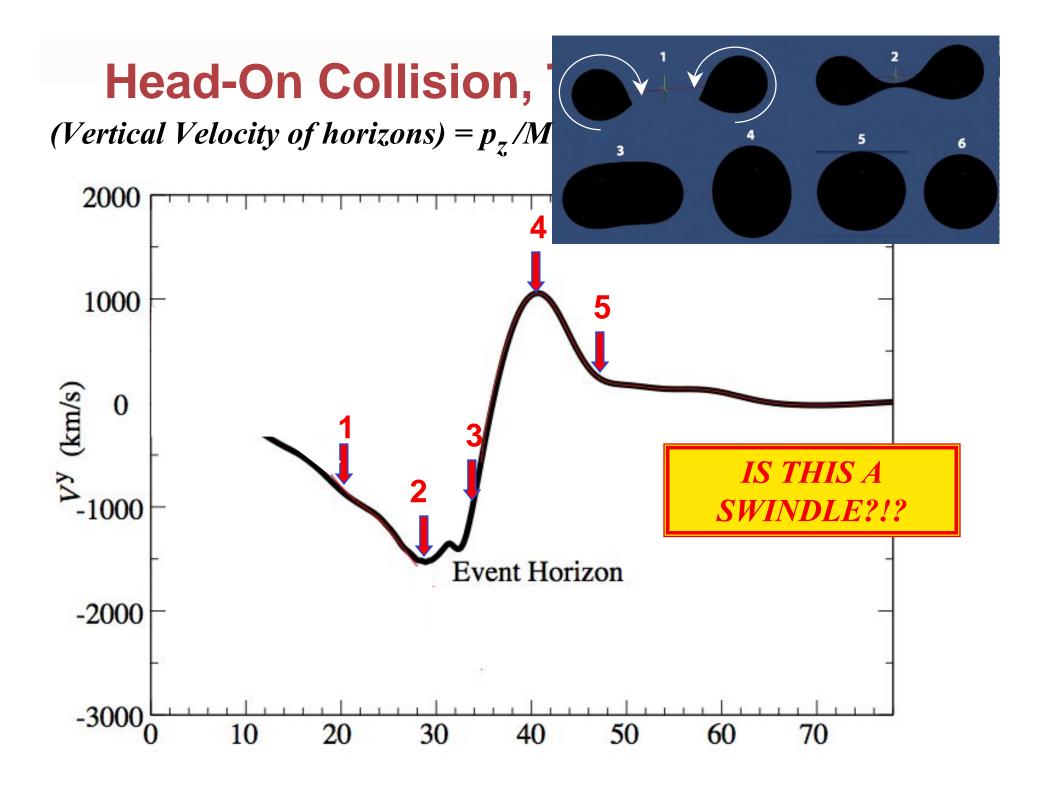
in Post-Newtonian approximation: Chandrasekhar or harmonic gauge



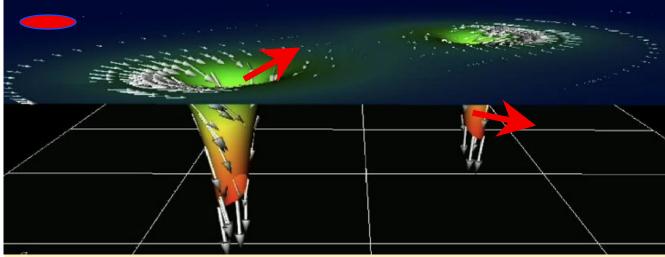
Head-On Collision, Transverse Spin

Geoffrey Lovelace, Mark Scheel, Michael Cohen, Jeff Kaplan [Cornell/Caltech]





How Define Momentum in Curved Spacetime?



Result of transport depends on path

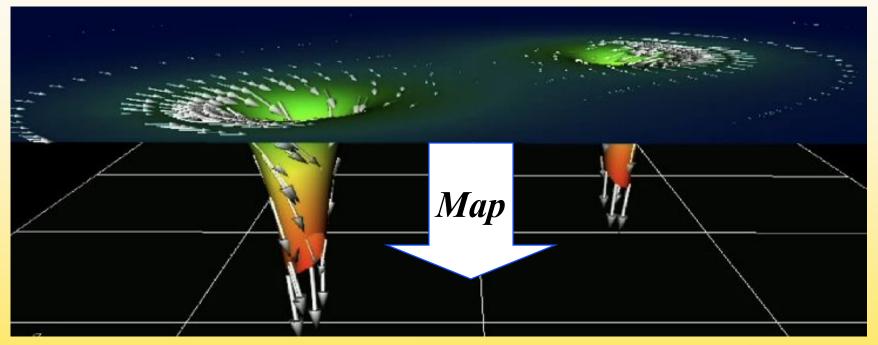
Must transport momentum vectors to a common location before adding them

Momentum conservation arises from translation invariance of spacetime.

BH/BH spacetime has none.

How Define Momentum in Curved Spacetime?

Rewrite General Relativity Theory as a Nonlinear Field Theory in Flat Spacetime

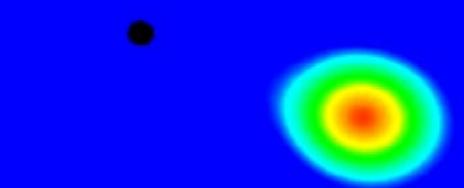


Landau & Lifshitz, Classical Theory of Fields

Mapping is not unique. - Momentum "gauge dependent"

-- but not "very" gauge dependent in this case --

G. Lovelace et al, arXiv:0907.0869



Fundamental Physics Issues

The dynamics of spacetime near generic singularities

- » Penrose-Hawking Singularity Theorems (1964 72)
- » Belinsky-Khalatnikov-Lifshitz (BKL) singularity is it truly generic?

Has been confirmed generic in NR simulations, by David Garfinkle

Future Challenge: How does singularity evolve as hole ages?

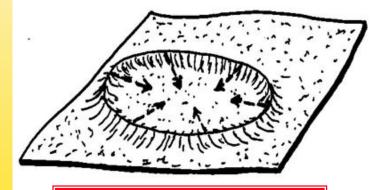
Chaotic pattern of stretch and squeeze

Fundamental Physics Issues

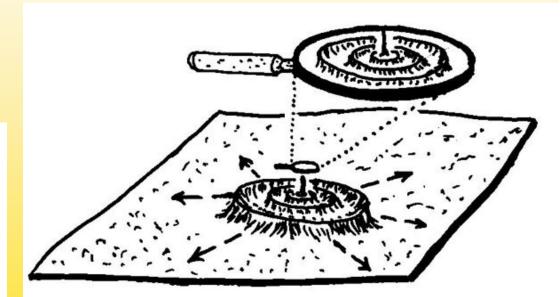
• Cosmic Censorship Conjecture [Penrose 1968]

» All singularities (except the big bang) are hidden inside black holes.

Numerical Simulations Matt Choptuik ~ 1994 -(U Texas -> UBC)



Imploding Scalar Waves

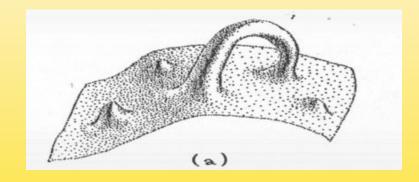


Critical behavior & scaling

Naked Singularity

Fundamental Physics Issues: future

- Topological Censorship Theorem [John Friedman et al 1993]:
 - » If the stress-energy tensor always satisfies the null energy condition (NEC), $T_{ab} k^a k^b ≥ 0$ for all null k^a , then information can never travel through a wormhole.
- BUT: null energy condition can be violated, e.g. in Casimir effect and in "squeezed vacuum"

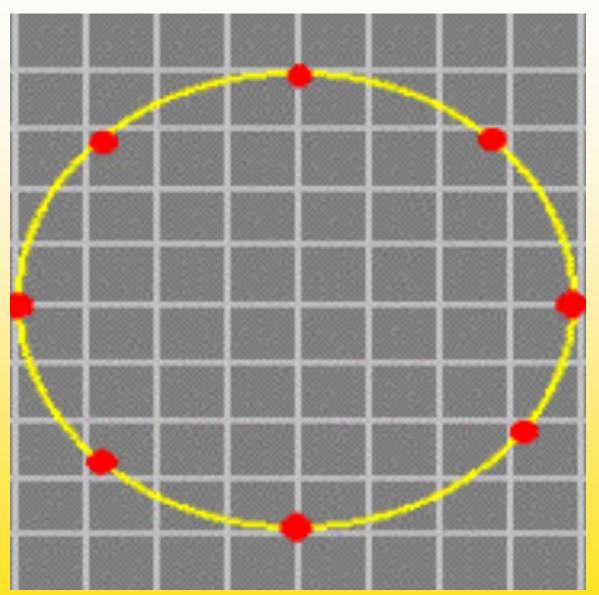


•How much violation of NEC is needed to permit travel through a wormhole? [NR]. How much is allowed? [QFT]

•What is the dynamics of the singularity that prevents travel? [NR]

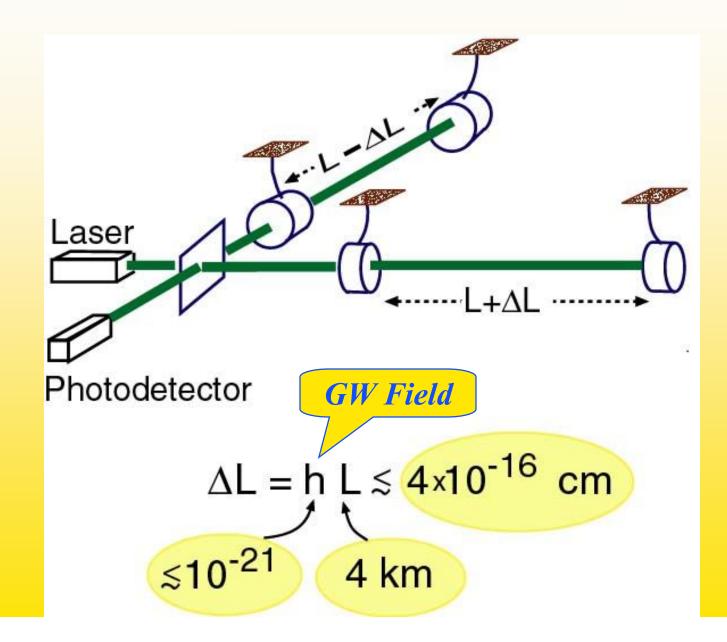
Part 2 Gravitational Wave Observations Probe the Warped Side of Universe

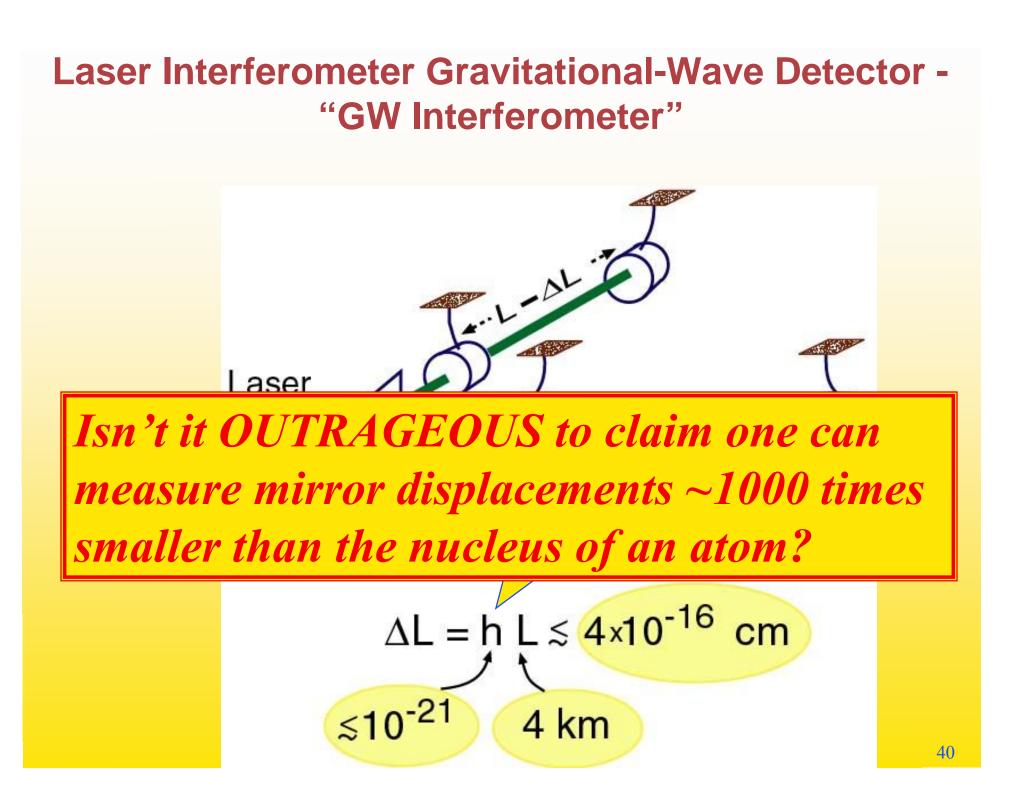
Motivation

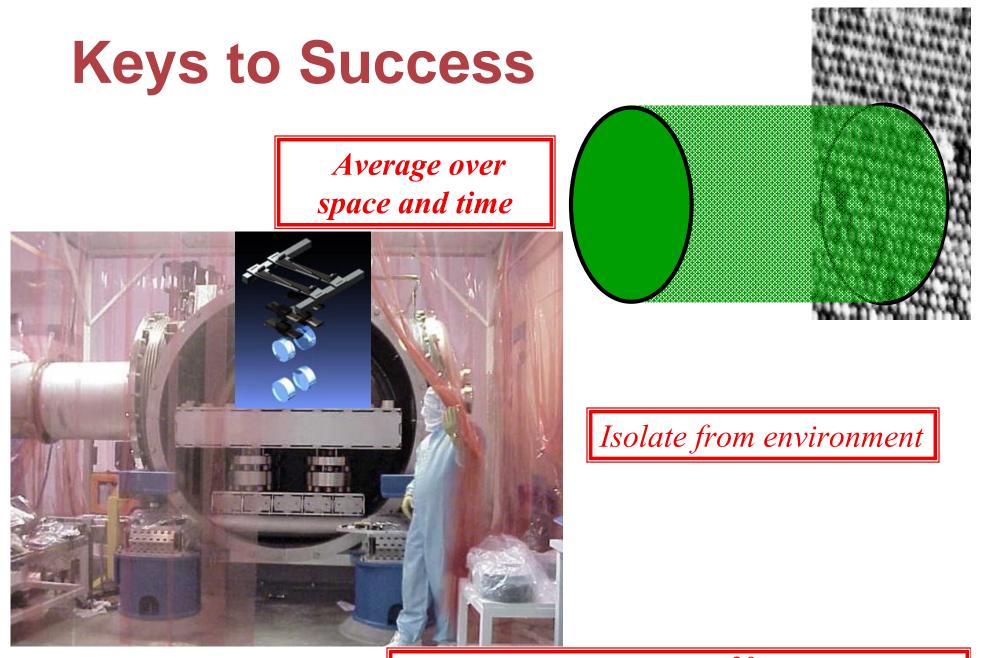


 $\Delta L/L = h(t)$

Laser Interferometer Gravitational-Wave Detector -"GW Interferometer"







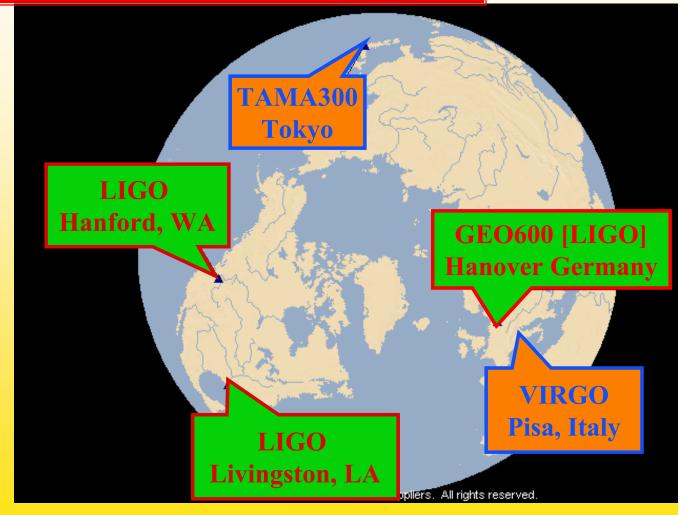
Use lots of photons: $\sim 10^{20}$ in 0.01 second

Earth-Based GW Interferometers Small holes in distant galaxies:

~10 to 100 Msun . ~ 100 km size

Network Required for:

- » Detection Confidence
- » Waveform Extraction
- » Direction by Triangulation



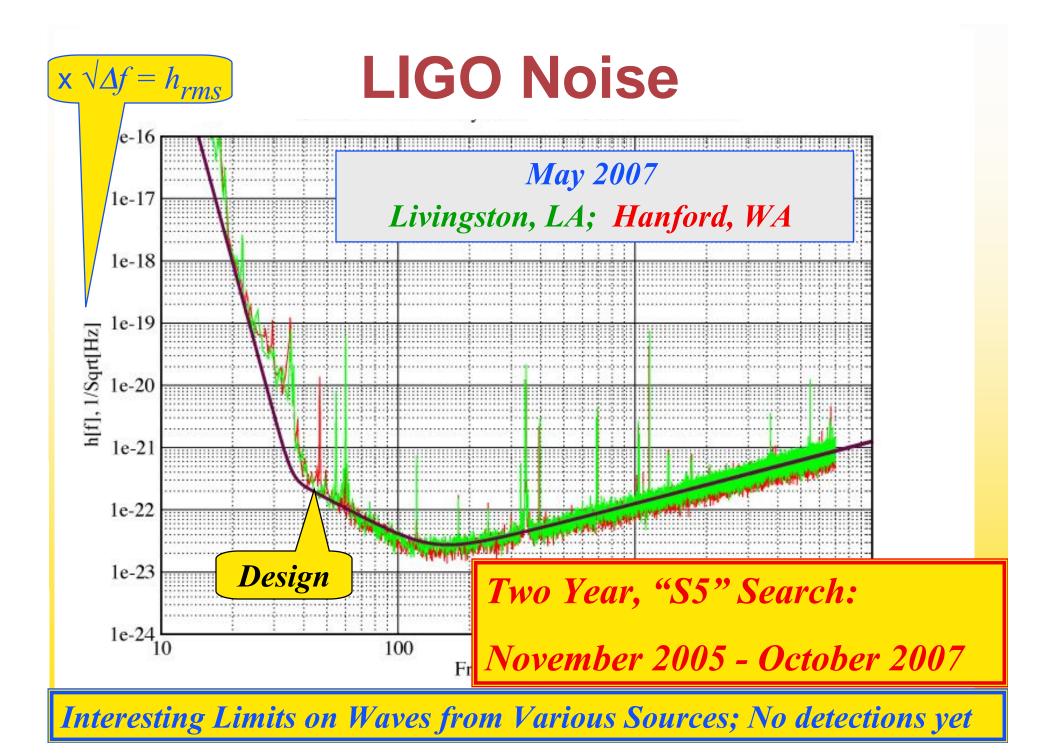
LIGO: Laser Interferometer Gravitational Wave Observatory Began as MIT/Caltech collaboration [Weiss; Drever, Kip] Now: Collaboration of ~500 scientists at ~50 institutions in 8 nations [J. Marx, Director; D. Rietze, Spokesman]





USA, UK, Germany, Australia, India, Japan, Russia, Spain





A few S5 Results from ~1/2 of Data

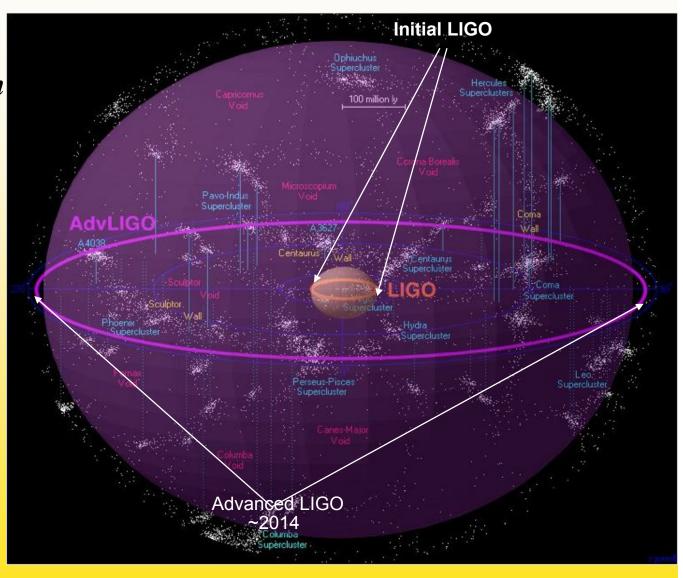
- BH/BH Binaries with M_{tot} < 35M_{sun}: <1/860 yrs in MWEG
- GRB070201 (coincident with Andromeda) is not a NS/NS or NS/BH in Andromeda
- Targeted Pulsar Search
 - » Crab pulsar: < 7% of spindown energy goes to GWs
- Stochastic Background: Ω< 7 x 10⁻⁶ in 41-178 Hz band (Bayesian 90% confidence)

Future Interferometers in LIGO

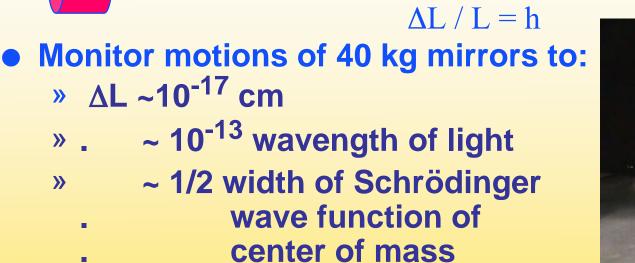
Initial LIGO: BH/BH ~300 million light years -≤ 1 BHBH / 10 yrs

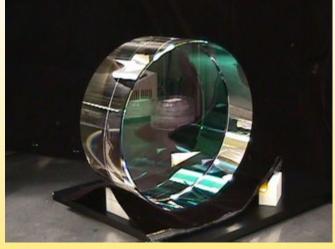
Enhanced: 2009-10 ~600 *million lt yrs* ≤ 1 *BHBH / yr*

Advanced: 2014-... ~5 billion lt yrs ~1 BHBH/day - mo



Advanced LIGO Interferometers The Experimental Challenge





48

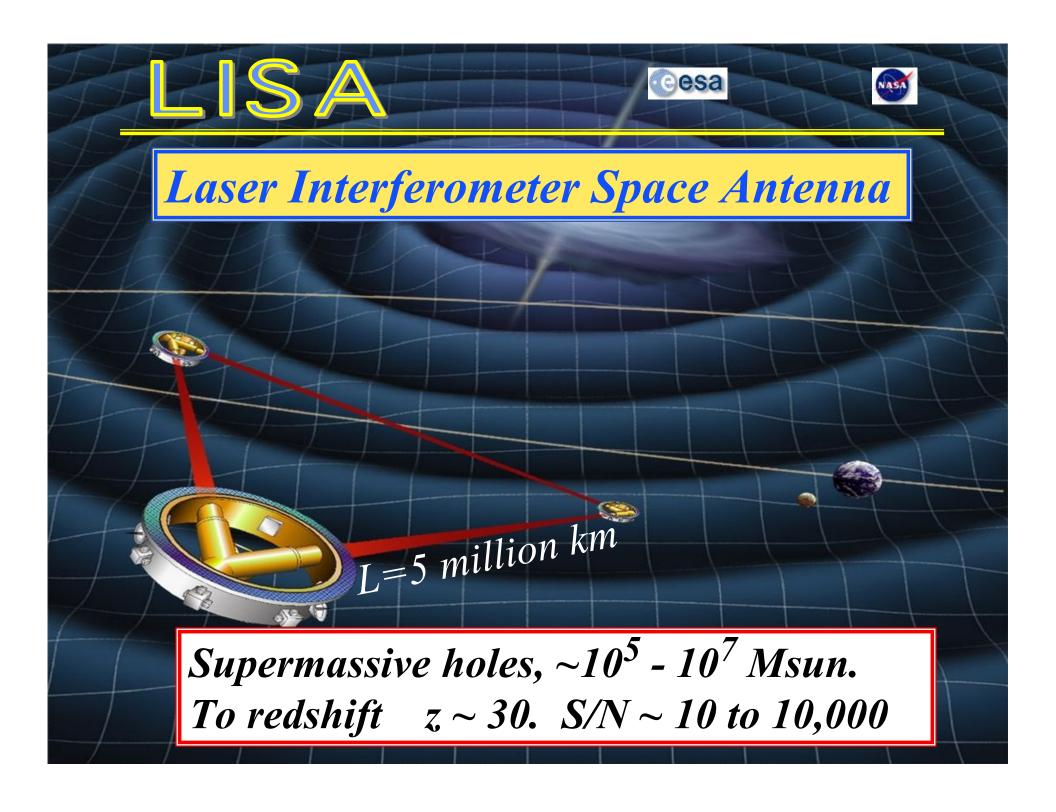
For the first time humans will see human-sized objects behave quantum mechanically!

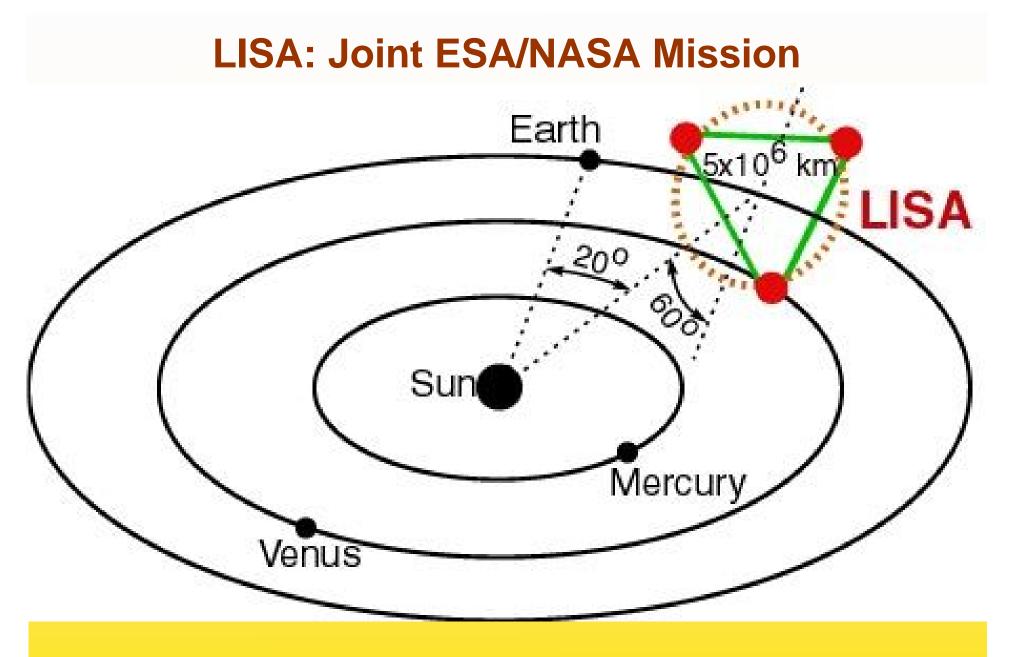
Quantum Nondemolition (QND) Technology to deal with this

Oct 2 Lorentz Lecture: Advanced LIGO - 2 modes of operation

• GW searches: insensitive to quantum state of mirrors

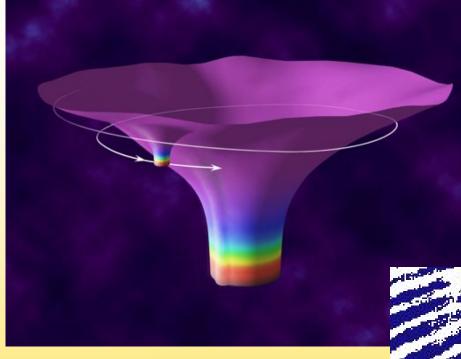
Macroscopic QM experiments: maximally sensitive to quantum state





• Launch: about 2018 or later

Mapping a Quiescent Black Hole







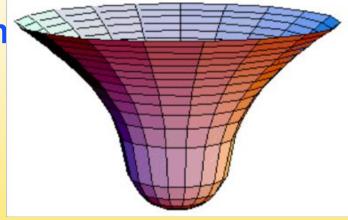
Some Numbers for LISA 5 million km = 20 light sec **Final Year:** 100,000 orbits with *Circumference* < 4 x (*Horizon circumference*) 3 billion light yrs 10 Msun 1 million Msun **L+∆** L h~10-20 **L=5** million km $\Delta L = 10^{-8} \, {\rm cm}$

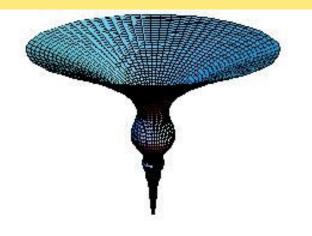
What if the Map is Not that of a Black Hole?

May have discovered a new type of "inhabitant" of dark side of the universe. Two long-shot possibilities:

 Dense objects made from cold, dark matter
 » (Dark ``Stars!!)
 » e.g. boson stars

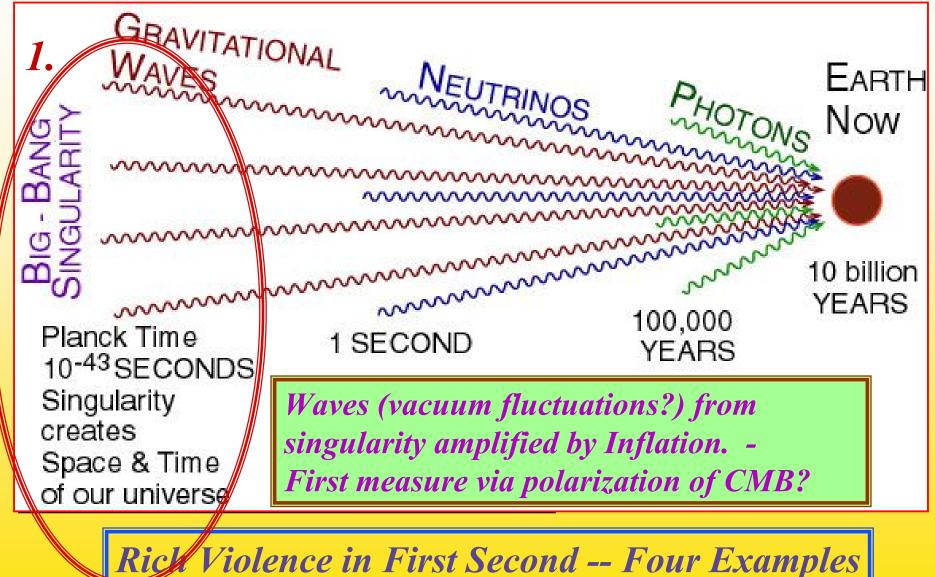
Naked Singularities





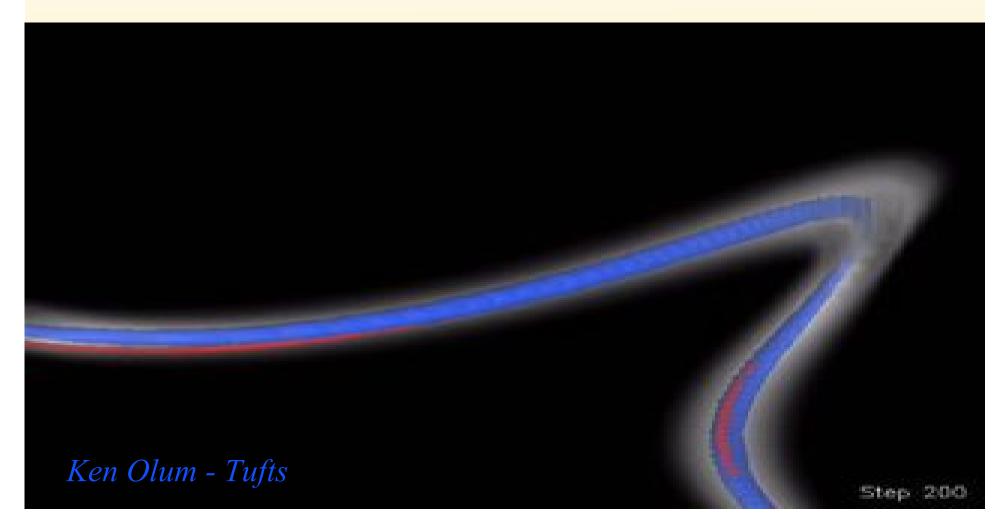
Over the Next 40 Years

Probe the Initial Second of Universe's Life



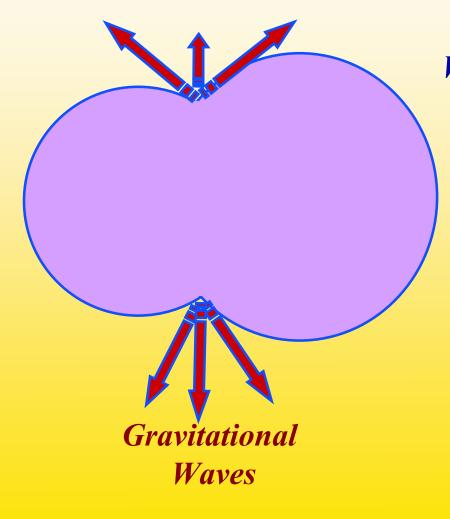
2. Cosmic Strings

- Inflation enlarges some superstrings to cosmic size
- Kinks, cusps and waves on cosmic strings produce gravitational waves



3. Birth of Fundamental Forces

- At age ~ 10^{-12} seconds [kT~ 1 TeV]:
 - » Phase transition: Electroweak force -> EM + Weak



Waves are in LISA's domain

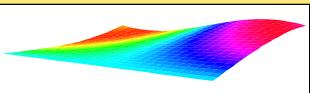
LIGO: Probe physics at age ~ 10⁻²² seconds [kT~10⁵ TeV]

4. Our 3-D Universe as a "Brane" in Higher Dimensional Bulk

- May have formed wrinkled
- As universe expanded, adjacent regions discovered the wrinkle between them



 Wrinkle began vibrating -- producing gravitational waves - brane smoothed out



Example of the kind of surprise gravitational-waves may bring us

Conclusions

Numerical Relativity and Gravitational Wave Observations are on the threshold of producing a revolution in our knowledge of the Warped Side of our Universe

